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imagery analysis report.

lonospheric Research At Dushambe, USSR (S)



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antennas. The heater was first observed under construction in May 1976 and was completed in August 1980. It consists of 25 masts, each down the center of each row. The outer rows of masts are meters apart. The masts support a series of symmetrical dipole antennas arranged in an undetermined configuration. When the heater is turned on, there is an interaction between the beam of radio waves produced by the heater and the electrons in the ionosphere at the point where the beam is directed. This interaction causes the ionospheric electrons to absorb energy from the beam and, as a result, to be pushed away from the point of disturbance. This creates a hole in the ionosphere. A communications signal is then beamed into this hole and propagated for long distances.1

7. (S/WN) The ten Yagi antennas and the FORK REST antenna were added to the facility between 1976 and 1981. They are probably used in frequency mixing experiments in conjunction with the heater. The SIDE NET radar was added in July 1976 and may be used to track meteor trails through the ionosphere, perhaps as an adjunct to the special ionospheric research station discussed below.

Special Ionospheric Research Station

- 8. (S/WN) The special ionospheric research station (Figure 3) is in the eastern portion of the Dushanbe facility. It consists of 4 square antenna arrays, a curtain array, 2 Yagi antennas, a tower, 5 control buildings, and 12 housing/support buildings. The station is involved in meteor radio electronics research. It was here that the drift of meteor trails and ionospheric irregularities was measured during 1968, 1969, and 1972 by a radio method.3 As meteors burn up in the ionosphere they leave ionized trails. The winds in the ionosphere can be measured by tracking these ionized trails with radio echoes. Ionospheric irregularities can be created naturally, by sunspot activity for example, or they can be man-made through various means, including the use of ionospheric heaters such as the Dushanbe heater.
- 9. (S/WN) Each of the four square antenna arrays at the station consists of five towers (four corner towers and a central tower), with a control building at the base of each of the central towers. Three of the arrays are identical in size while the fourth is slightly larger. The three identical arrays are 53 meters square and have 22-meter-tall towers. The fourth array is 65 meters square with 26-meter-high corner

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- 1 -**SECRET** towers and a 54-meter-high central tower. An additional tower, 26 meters high, with a control building at its base, is near the center of the four antenna arrays. A Yagi antenna has been on one of the corner towers of the easternmost square antenna array since May 1976. A second Yagi antenna is near the curtain array. Yagi antennas have also been observed on the central towers in each square antenna array at various times. A Soviet source described the antennas used in measuring the winds in the lower ionosphere using meteor trail drift as "four pairs of receiving and transmitting antennas oriented in the N/S and E/W directions".4 The four antenna arrays at the special ionospheric station are oriented exactly on a north/south and east/west axis and were the only ones present from 1968 through 1969, and in 1972. They are probably the ones used to track meteor trail drift in the ionosphere. The function of the Yagi antennas has not been determined.

10. (S/WN) The curtain array consists of two 43-meter-high towers, 35 meters apart, each with five horizontal cross arms. Each cross arm is long, and the cross arms are antenna was added between August 1977 and June 1978. The function of this antenna has not been determined.

Visual and Photographic Observation Station and Astronomical Observatory

11. (S/WN) The visual and photographic observation station, designated station 1068 of the Soviet Astronomical Council, and the collocated Astronomical Observatory of the Academy of Sciences Tadzhik SSR5 (Figure 4) are in the western portion of the Dushanbe facility. Station 1068 consists of one large optical tracking building with a VAU (high-precision astronomical apparatus) camera, two triple-position optical-tracking buildings, five small, single-position optical-tracking buildings, two parabolic dish antennas (one covered by a movable building), and one mast with an unidentified antenna on top. There are also ten support buildings, one control building, three computer vans, and six generator vans. The 35-ton VAU camera is an improved device for optical observation of satellites and is similar to the US Baker-Nunn camera. The collocated astronomical observatory consists of three domes housing optical telescopes used in the study of celestial mechanics and phenomena.

12. (S/WN) Station 1068 is devoted primarily to optical and photographic tracking of satellites and secondarily to ionospheric research. The optical tracking positions contain cameras which photograph satellite tracks against a star background. This is usually done just after sunset when the sky is dark but the satellites are still illuminated by sunlight. The same cameras are used in ionospheric research projects, such as photographing the behavior of barium clouds that have been injected into the ionosphere by rockets.7

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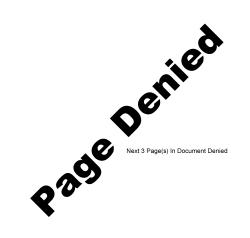
(S/WN) All relevant satellite imagery acquired from January 1970 through December 1981 was used in the preparation of this report.

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